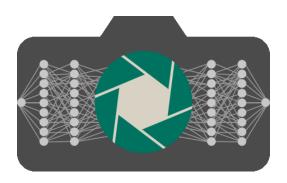
WashU Vision & Learning Group

https://vlg.seas.wustl.edu/



PI: Ayan Chakrabarti (ayan@wustl.edu)

An Overview of Computer Vision & Computational Photography

CSE 591: Sep 19, 2018

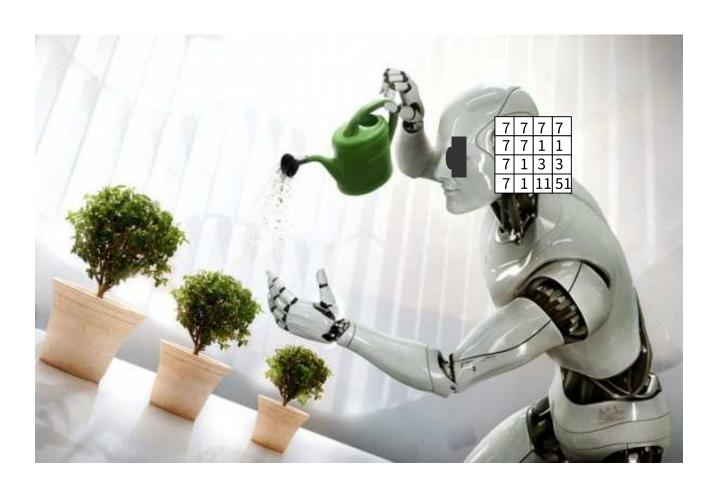
What is Computer Vision?

Computational Systems to make sense of the physical world by looking at images and videos

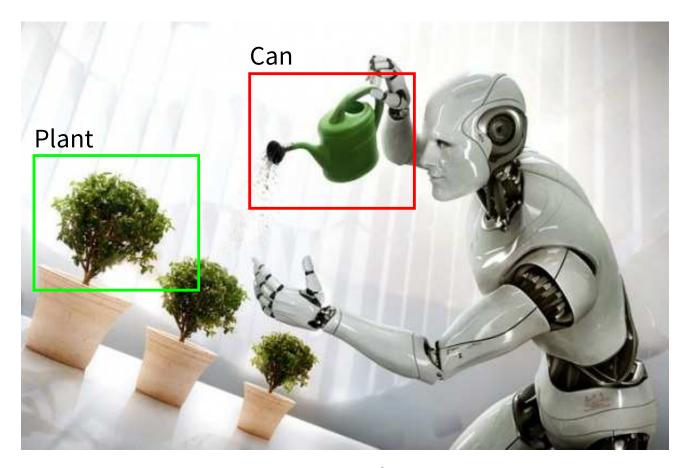


[credit: http://www.blutsbrueder-design.com]

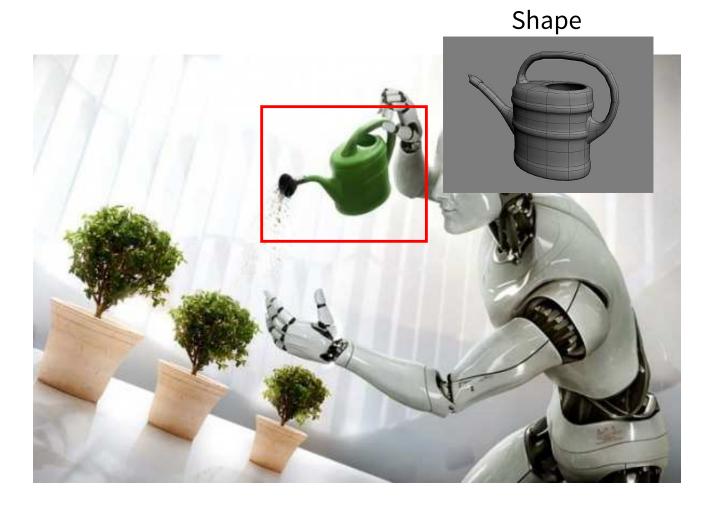






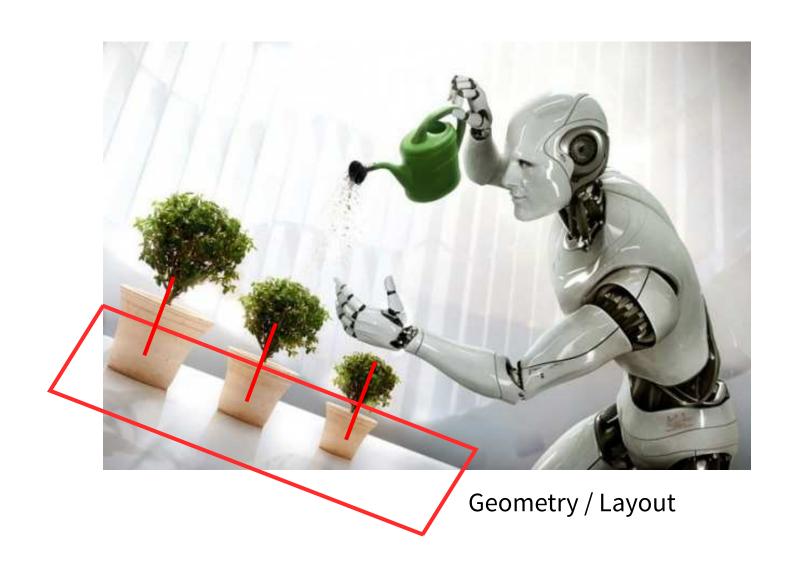


Recognize Objects



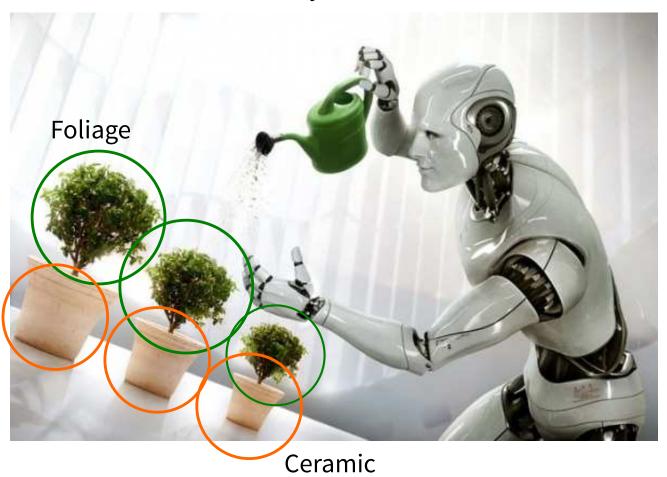


Classify Scene





Identify Materials



Surface Properties



Computer Vision

• Develop Algorithms that extract a description of the world from images

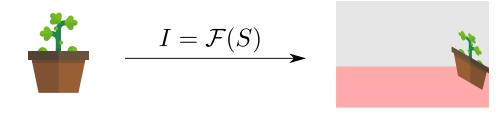
Computational Photography

• Think of modified cameras and acquisition setups that make this extraction easier

Broad Overview of (many a) Vision Algorithm

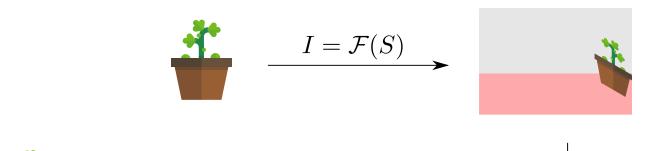
Broad Overview of (many a) Vision Algorithm

1. Understand the Image Formation Model: Scene to Image



Broad Overview of (many a) Vision Algorithm

1. Understand the Image Formation Model: Scene to Image





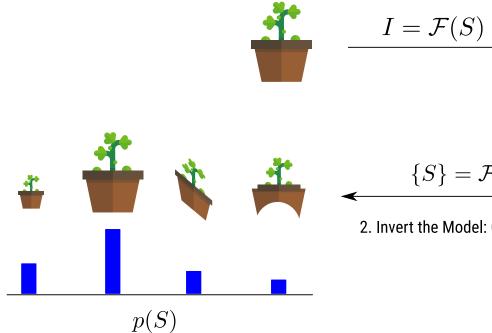


$$\{S\} = \mathcal{F}^{-1}(I)$$

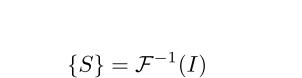
2. Invert the Model: Gives us Multiple Physically Feasible Solutions

Broad Overview of (many a) Vision Algorithm

1. Understand the Image Formation Model: Scene to Image



3. Learn What Natural Scenes Look Like: Use to select likely scene among those that are feasible

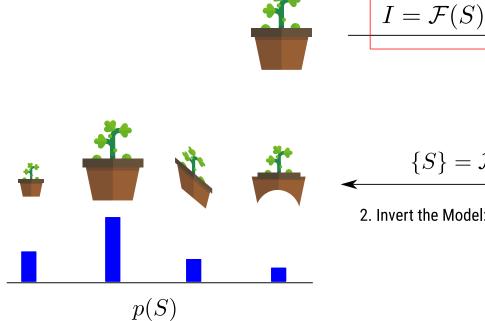


2. Invert the Model: Gives us Multiple Physically Feasible Solutions

Broad Overview of (many a) Vision Algorithm

1. Understand the Image Formation Model: Scene to Image

4. Computational Photography: Modify the Image Formation Model to make measurements more informative



$$\{S\} = \mathcal{F}^{-1}(I)$$

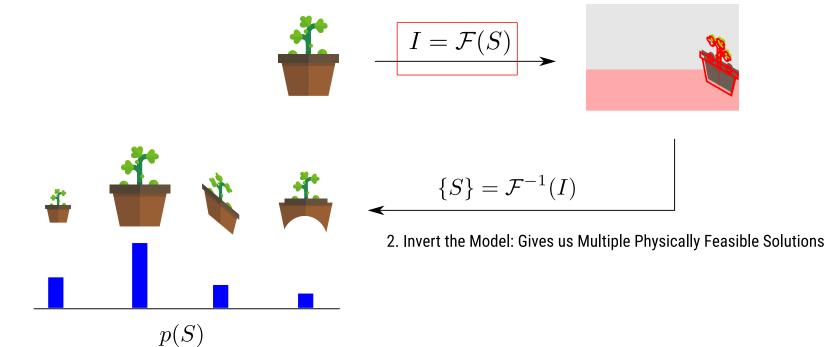
2. Invert the Model: Gives us Multiple Physically Feasible Solutions

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Broad Overview of (many a) Vision Algorithm

1. Understand the Image Formation Model: Scene to Image

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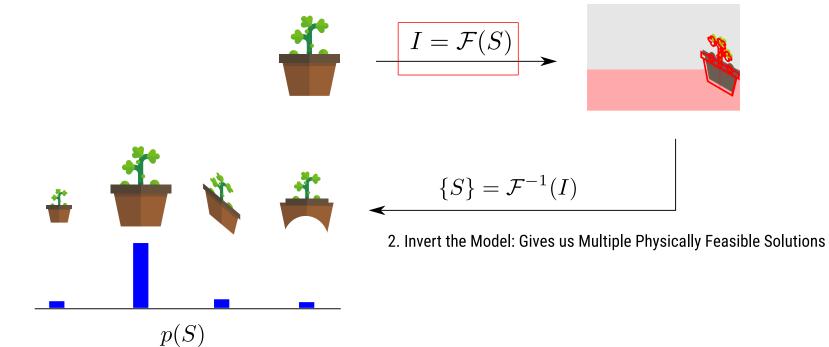


3. Learn What Natural Scenes Look Like: Use to select likely scene among those that are feasible

Broad Overview of (many a) Vision Algorithm

1. Understand the Image Formation Model: Scene to Image

4. Computational Photography: Modify the Image Formation Model to make measurements more informative



3. Learn What Natural Scenes Look Like: Use to select likely scene among those that are feasible

REAL-WORLD IMPACT

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Bronze











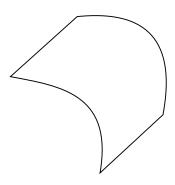


WHAT DOES VISION RESEARCH LOOK LIKE?

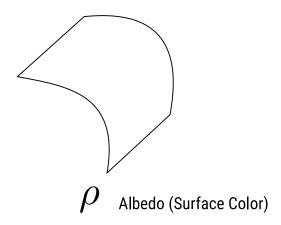
As a Grad Student working on a problem, you will have to:

- Understand the physics, geometry, optics, etc. of the setup.
- Understand to what degree the image formation process is invertible, characterize the ambiguity.
- Figure out how to the statistics of natural images could resolve this ambiguity.
- Use this to choose a model / architecture.
- Figure out how to train / learn parameters of this model.
- Develop an algorithm to use this model for actual inference.
- Make sure this is efficient and practical.

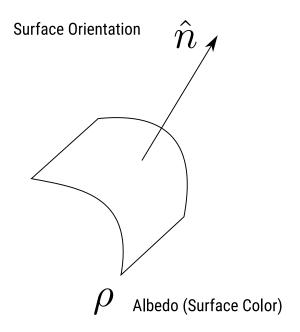
Photometric Stereo



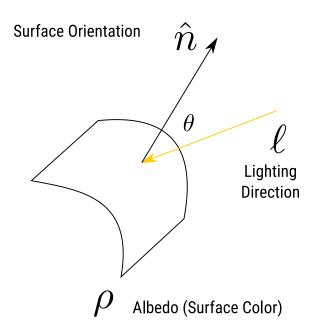
Photometric Stereo



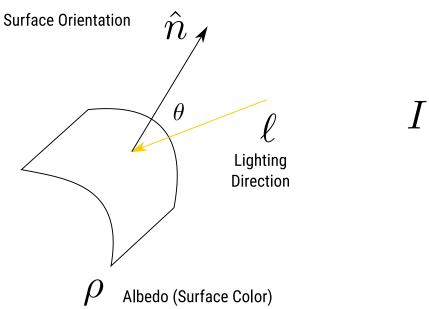
Photometric Stereo



Photometric Stereo

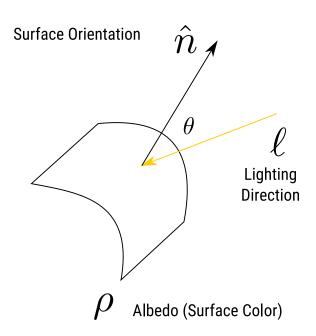


Photometric Stereo



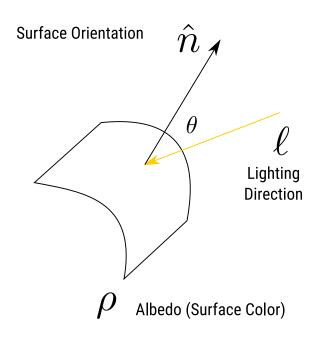
$$I = \rho \cos \theta$$

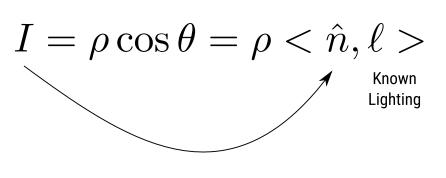
Photometric Stereo



$$I = \rho \cos \theta = \rho < \hat{n}, \ell >$$

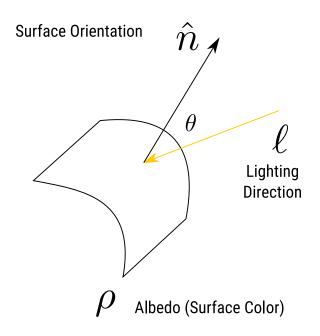
Photometric Stereo

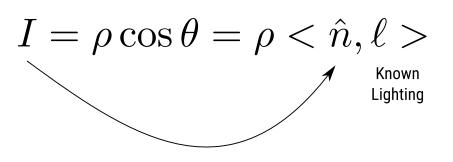




Photometric Stereo

Use the fact that intensity depends on relative angle between surface normal and light source.

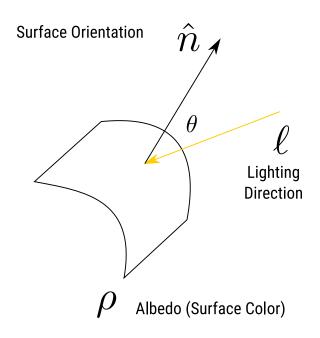


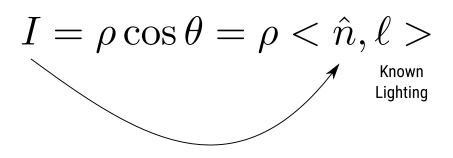


One Observation: Three Unknowns

Photometric Stereo

Use the fact that intensity depends on relative angle between surface normal and light source.





One Observation: Three Unknowns

Take multiple images with different lighting

Photometric Stereo



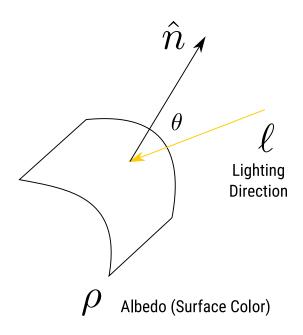
Photometric Stereo



Great, but requires you to take multiple images. What if the object is moving?

RGB Photometric Stereo

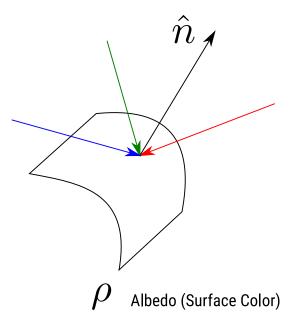
Take three shots in one: use an RGB Camera



$$I = \rho < \hat{n}, \ell >$$

RGB Photometric Stereo

Take three shots in one: use an RGB Camera

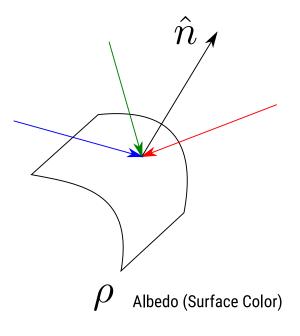


$$\begin{split} I_{_{\mathrm{R}}} &= \rho_{_{\!\mathrm{R}}} < \hat{n}, \ell_{_{\mathrm{R}}} > \\ I_{_{\!\mathrm{G}}} &= \rho_{_{\!\mathrm{G}}} < \hat{n}, \ell_{_{\!\mathrm{G}}} > \\ I_{_{\!\mathrm{B}}} &= \rho_{_{\!\mathrm{B}}} < \hat{n}, \ell_{_{\!\mathrm{B}}} > \end{split}$$

RGB Photometric Stereo

Take three shots in one: use an RGB Camera

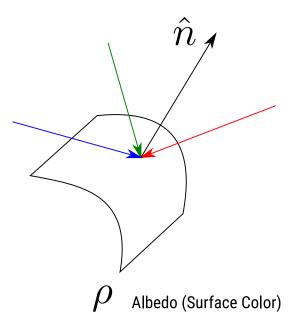
But now we have extra unknowns for surface color



$$\begin{split} I_{_{\mathrm{R}}} &= \rho_{_{\!\mathrm{R}}} < \hat{n}, \ell_{_{\mathrm{R}}} > \\ I_{_{\!\mathrm{G}}} &= \rho_{_{\!\mathrm{G}}} < \hat{n}, \ell_{_{\!\mathrm{G}}} > \\ I_{_{\!\mathrm{B}}} &= \rho_{_{\!\mathrm{B}}} < \hat{n}, \ell_{_{\!\mathrm{B}}} > \end{split}$$

RGB Photometric Stereo

Take three shots in one: use an RGB Camera



But now we have extra unknowns for surface color

Solution 1: Measure albedo separately (assuming it's constant)

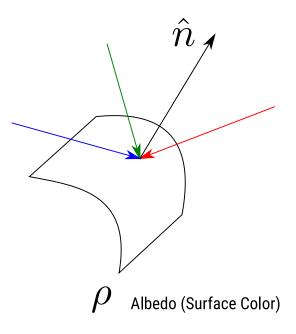
$$I_{_{\mathrm{R}}}=\rho_{_{\!\mathrm{R}}}<\hat{n},\ell_{_{\mathrm{R}}}>$$

$$I_{\rm g} = \rho_{\rm g} < \hat{n}, \ell_{\rm g} >$$

$$\begin{split} I_{_{\mathrm{R}}} &= \rho_{_{\!\mathrm{R}}} < \hat{n}, \ell_{_{\mathrm{R}}} > \\ I_{_{\!\mathrm{G}}} &= \rho_{_{\!\mathrm{G}}} < \hat{n}, \ell_{_{\!\mathrm{G}}} > \\ I_{_{\!\mathrm{B}}} &= \rho_{_{\!\mathrm{B}}} < \hat{n}, \ell_{_{\!\mathrm{B}}} > \end{split}$$

RGB Photometric Stereo

Take three shots in one: use an RGB Camera



But now we have extra unknowns for surface color

Solution 1: Measure albedo separately (assuming it's constant)

Solution 2: Paint the object, so that albedo is known and constant

$$I_{_{\mathrm{R}}}=\rho_{_{\!\mathrm{R}}}<\hat{n},\ell_{_{\mathrm{R}}}>$$

$$I_{\rm g} = \rho_{\rm g} < \hat{n}, \ell_{\rm g} >$$

$$\begin{split} I_{_{\mathrm{R}}} &= \rho_{_{\!\mathrm{R}}} < \hat{n}, \ell_{_{\mathrm{R}}} > \\ I_{_{\!\mathrm{G}}} &= \rho_{_{\!\mathrm{G}}} < \hat{n}, \ell_{_{\!\mathrm{G}}} > \\ I_{_{\!\mathrm{B}}} &= \rho_{_{\!\mathrm{B}}} < \hat{n}, \ell_{_{\!\mathrm{B}}} > \end{split}$$

Single-image RGB Photometric Stereo With Spatially-varying Albedo

Ayan Chakrabarti TTI-Chicago Kalyan Sunkavalli Adobe Research

Abstract

We present a single-shot system to recover surface geometry of objects with spatially-varying albedos, from images captured under a calibrated RGB photometric stereo setup-with three light directions multiplexed across different color channels in the observed RGB image. Since the problem is ill-posed point-wise, we assume that the albedo map can be modeled as piece-wise constant with a restricted number of distinct albedo values. We show that under ideal conditions, the shape of a non-degenerate local constant albedo surface patch can theoretically be recovered exactly. Moreover, we present a practical and efficient algorithm that uses this model to robustly recover shape from real images. Our method first reasons about shape locally in a dense set of patches in the observed image, producing shape distributions for every patch. These local distributions are then combined to produce a single consistent surface normal map. We demonstrate the efficacy of the approach through experiments on both synthetic renderings as well as real captured images.

CHEZ W IN

from a single image of an object with unknown spatiallyvarying albedo under unknown natural lighting. Although impressive given the inherent ambiguities in the SFS setup, their recovered geometries are typically coarse due to the use of strong smoothness priors, and their inference algorithm is computationally expensive. This is true even when known lighting is provided as input to their algorithm, primarily because it is designed to handle arbitrary and potentially ambiguous *natural* illumination environments.

In this paper, we show that efficient and high-quality surface recovery from a single image is possible, when using a calibrated lighting environment that is specifically chosen to be directly informative about shape. Specifically, we use the RGB (or color) photometric stereo (RGB-PS) setup [3, 11, 14], where an object is illuminated by three monochromatic directional light sources, such that each of the red, green, and blue channels in the observed image is "lit" from a different direction. For natural lighting, directional diversity in color has been shown to be informative towards shape [10]. But the benefits of this lighting setup for shape recovery can be better understood by interpreting it as one that multiplexes the multiple images of classical PS into the different color channels of a single image.

Proposition 1. Given noiseless observed intensities v(p) at a set of locations $p \in \Omega$ on a diffuse surface patch known to have constant albedo, i.e., $\kappa(p) = \kappa_{\Omega}, \forall p \in \Omega$, the true surface normals $\{\hat{n}(p) : p \in \Omega\}$ and common albedo κ_{Ω} are uniquely determined, if:

- 1. All intensities v(p) are strictly positive.
- 2. The true surface is non-degenerate in the sense that the set $\{\hat{n}(p)\hat{n}(p)^T:p\in\Omega\}$, of outer-products of the true normal vectors, span the space Sym_3 of all 3×3 symmetric matrices.

Proof: Given κ_{Ω} and $\hat{n}(p)$ as the true patch albedo and normals, let κ'_{Ω} , $\hat{n}'(p)$ be a second solution pair that also explains the observed intensities v(p) in the patch Ω . Since the observed intensities are strictly positive, this implies that the albedos κ_{Ω} , κ'_{Ω} are strictly positive as well, and further that no point is in shadow under any of the lights, i.e. $L^T \hat{n}(p)$, $L^T \hat{n}'(p) > 0$, $\forall p \in \Omega$. Then, since L^T is invertible, we can write

$$\operatorname{diag}[\kappa_{\Omega}]L^{T}\hat{n}(p) = \operatorname{diag}[\kappa'_{\Omega}]L^{T}\hat{n}'(p)$$

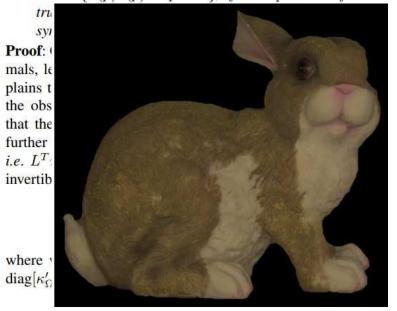
$$\Rightarrow \hat{n}'(p) = A\hat{n}(p), \ \forall p \in \Omega,$$
 (2)

where we define the matrix $A=L^{-T}RL^{T}$, with $R=\mathrm{diag}[\kappa'_{\Omega}]^{-1}\mathrm{diag}[\kappa_{\Omega}]$ being a diagonal matrix whose entries

Analyze ambiguities, and show that for "most" local regions, if the albedo is constant inside the region, we can recover its shape and albedo uniquely.

Proposition 1. Given noiseless observed intensities v(p) at a set of locations $p \in \Omega$ on a diffuse surface patch known to have constant albedo, i.e., $\kappa(p) = \kappa_{\Omega}, \forall p \in \Omega$, the true surface normals $\{\hat{n}(p) : p \in \Omega\}$ and common albedo κ_{Ω} are uniquely determined, if:

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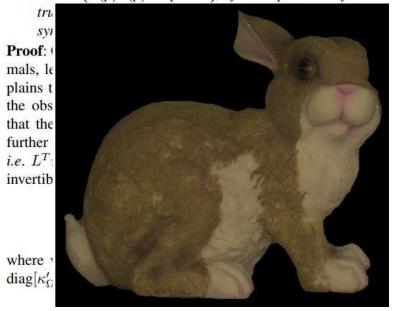


Analyze ambiguities, and show that for "most" local regions, if the albedo is constant inside the region, we can recover its shape and albedo uniquely.

But we don't know which patches are constant albedo, and which have boundaries in them.

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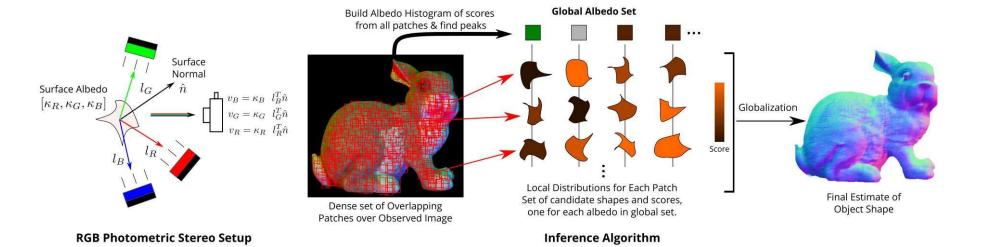
- 1. All intensities v(p) are strictly positive.
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Analyze ambiguities, and show that for "most" local regions, if the albedo is constant inside the region, we can recover its shape and albedo uniquely.

But we don't know which patches are constant albedo, and which have boundaries in them.

Also, uniqueness holds in idealized conditions. In reality, we'll have noise, 'non diffuse' reflection,





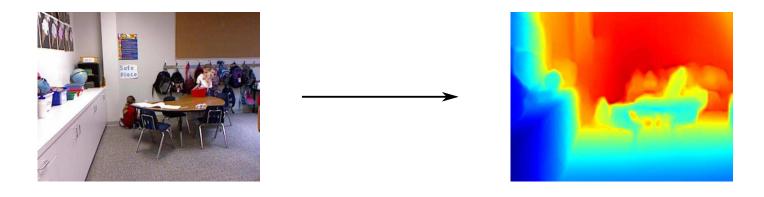






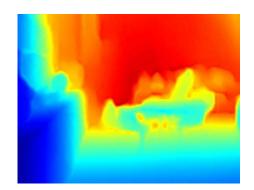






Depth from a Single Image





Depth from a Single Image

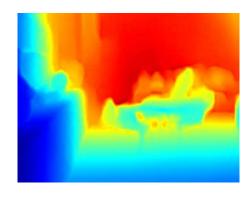
- No explicit geometric / optical cues.
- Must learn to map familiar patterns to depth.

Shading.

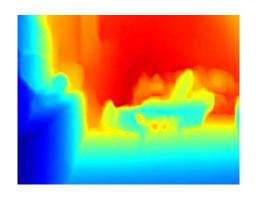
Contours & Boundaries.

Foreshortening of regular patterns.

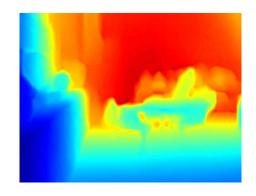
Scale of familiar objects.



Z(n)

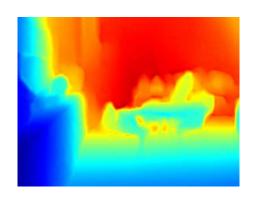


Z(n) Lots of numbers (200k for a 500x400 image)



Z(n)

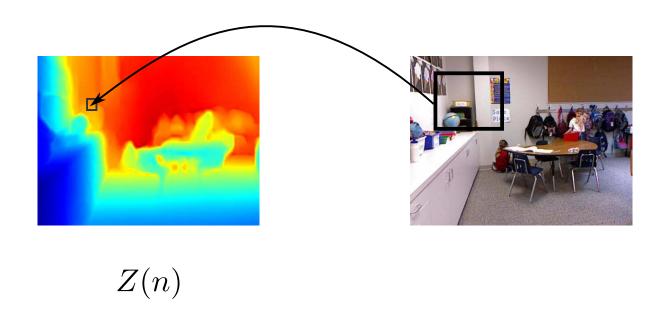
Estimate each Z(n) independently.

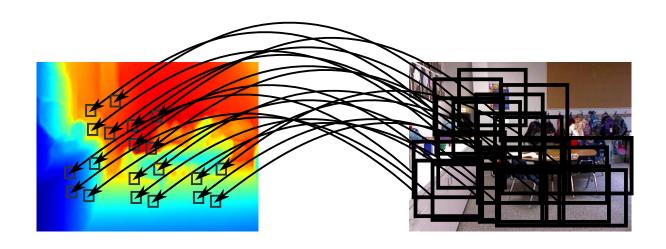




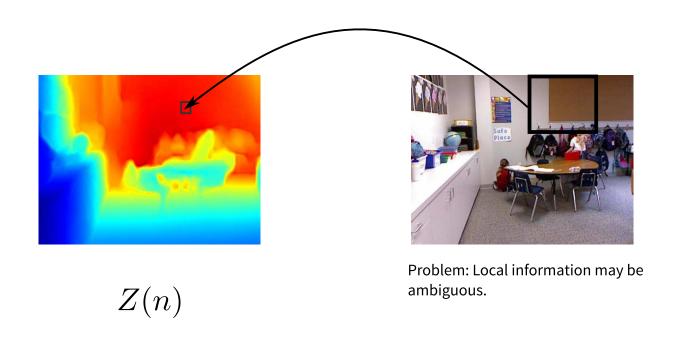


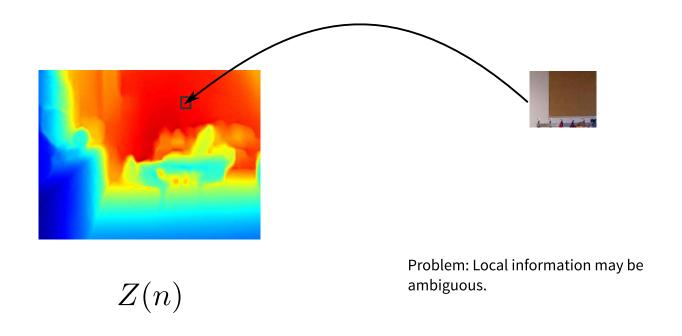
Estimate each Z(n) independently.

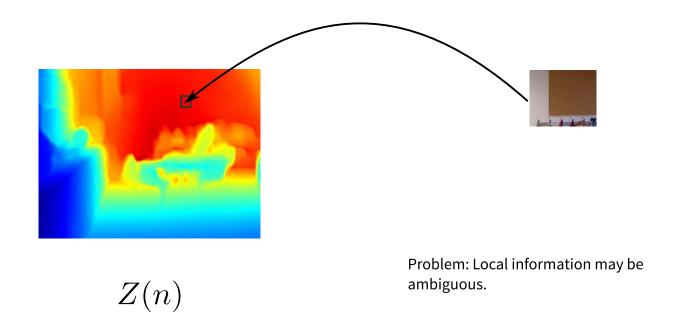




Z(n)

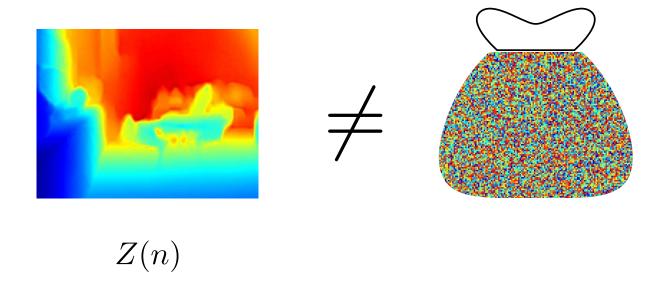




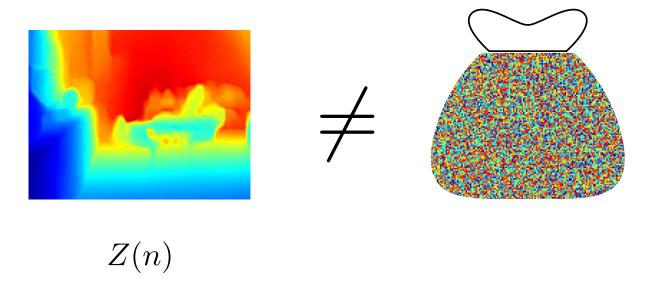


Estimate each Z(n) independently and *locally*.

Scene Maps have Structure



Scene Maps have Structure

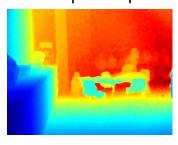


Build an algorithm that effectively extracts and exploits this structure

Scene Maps have Structure

Local Estimation: Mid-level Representation

Output Map

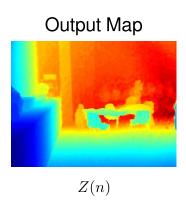


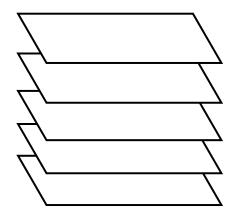
Z(n)



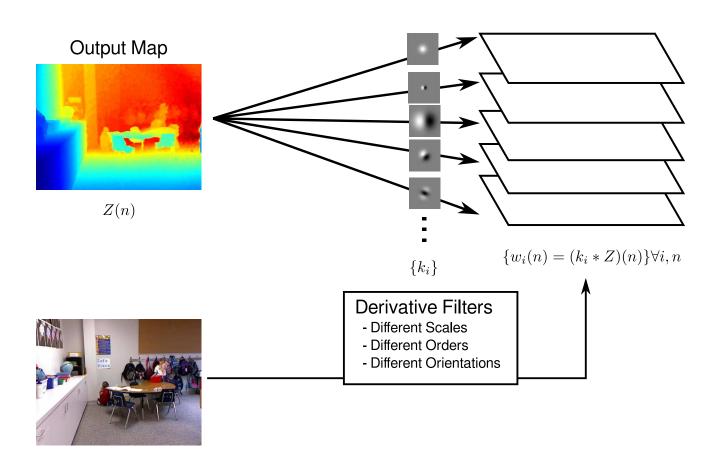


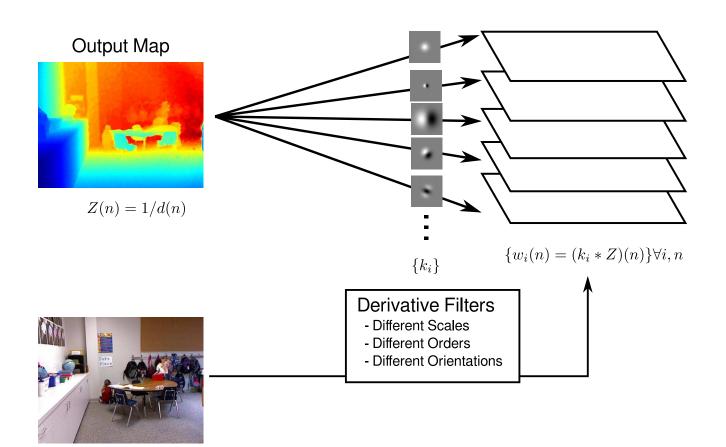
Input











Local Estimation: Mid-level Representation

Perspective Camera

$$d(x,y) \Rightarrow (dx,dy,d)$$

World 3D Co-ordinates

Plane Equation

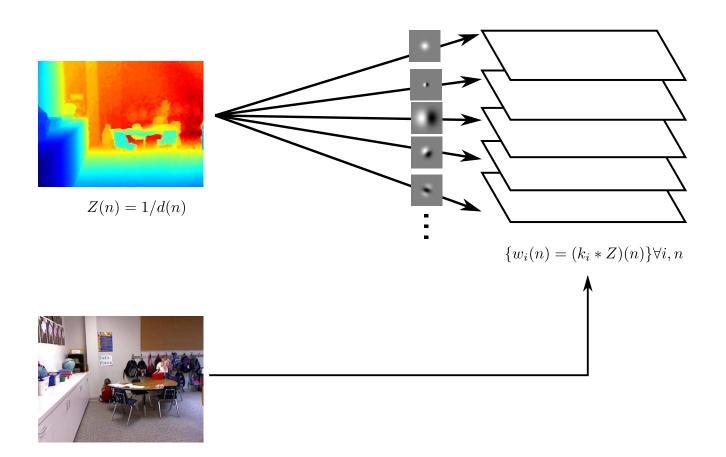
$$\frac{1}{d} = \alpha x + \beta y + \gamma$$

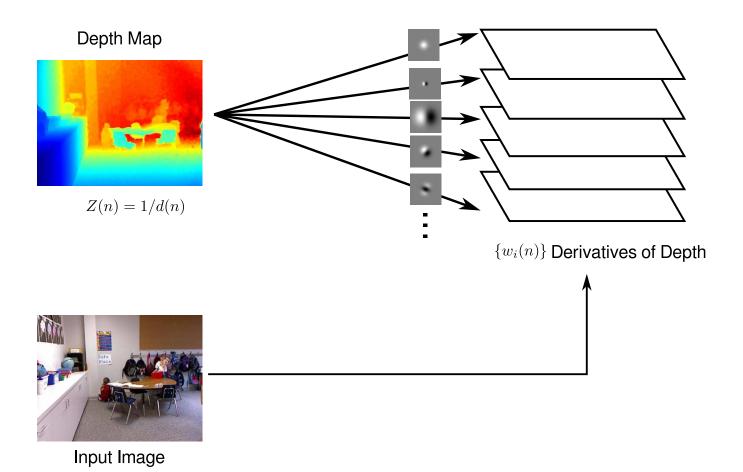
Zeroth Derivative = Absolute depth

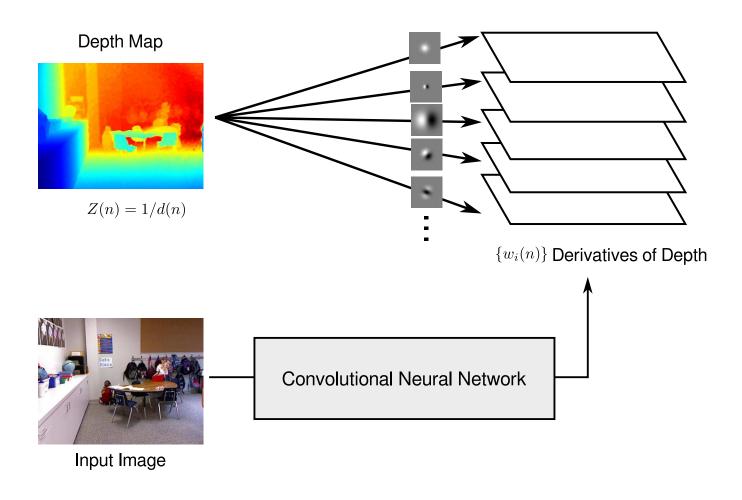
First Derivative = Surface orientation

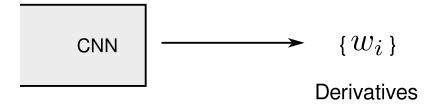
Second Derivative = 0: Planar

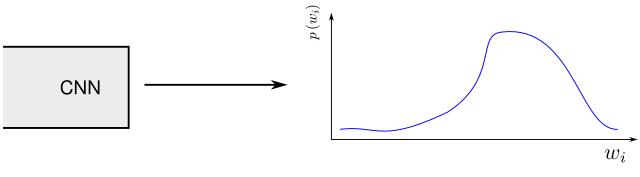
Curvature, contours.



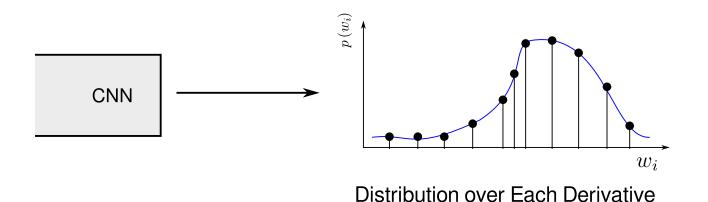


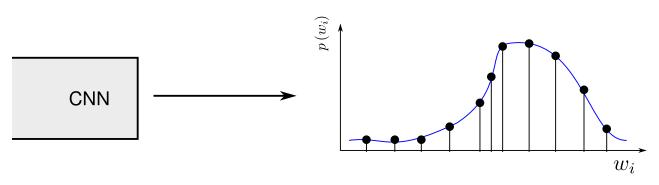






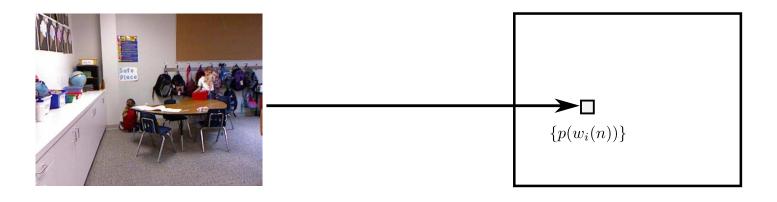
Distribution over Each Derivative



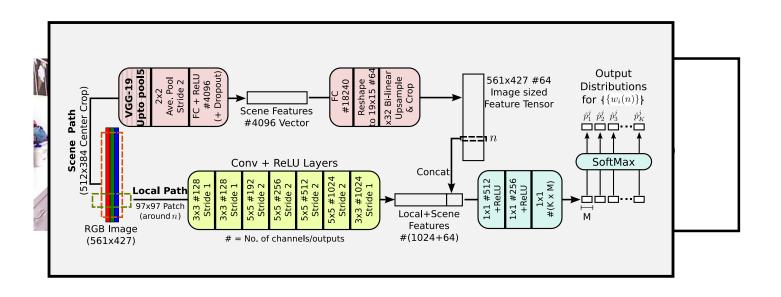


Distribution over Each Derivative

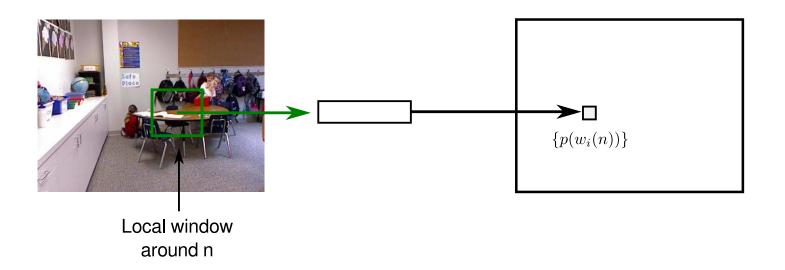
CNN



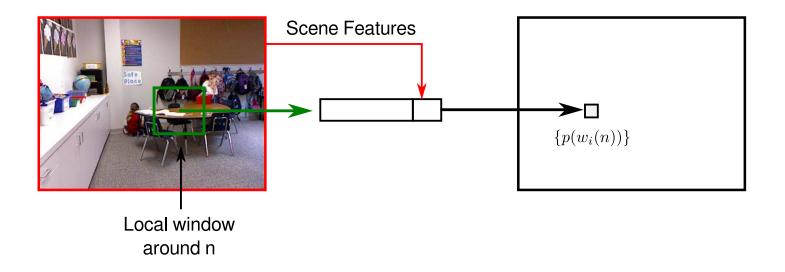
CNN

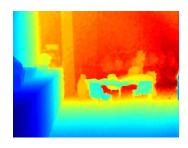


CNN



CNN

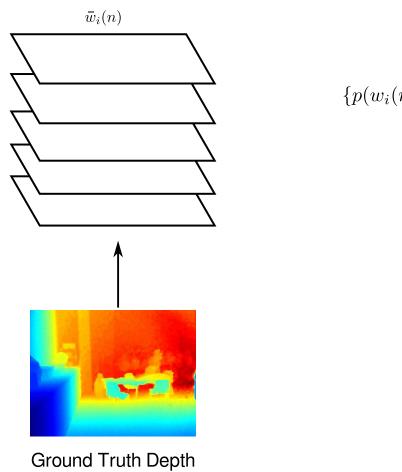


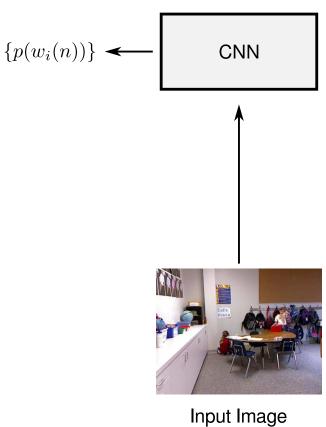


Ground Truth Depth

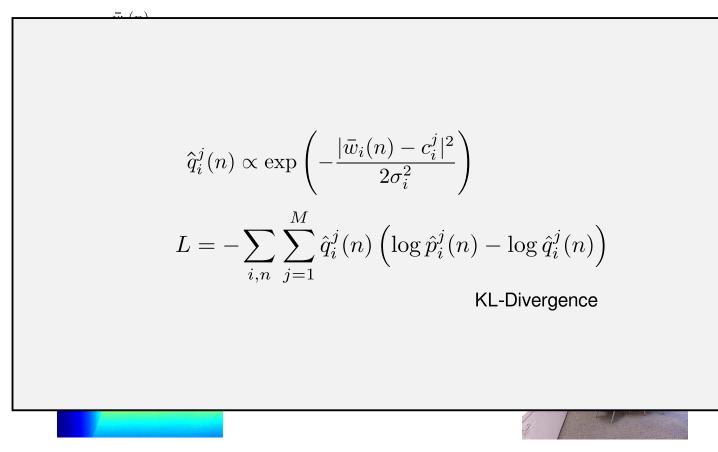


Input Image



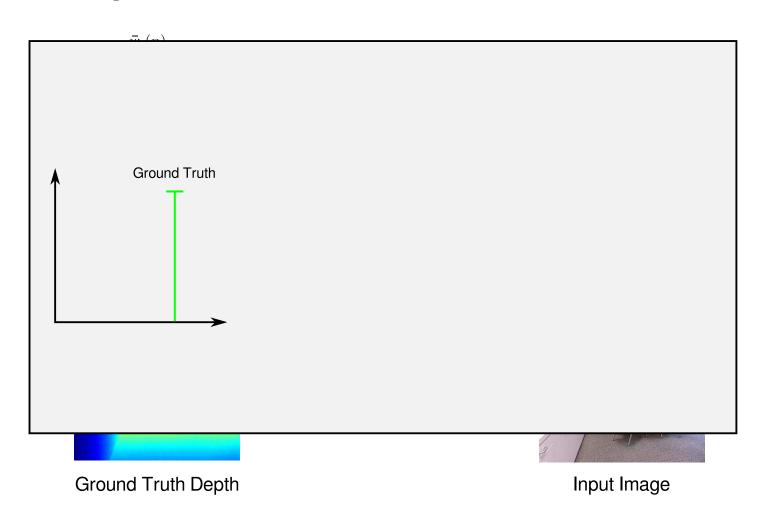


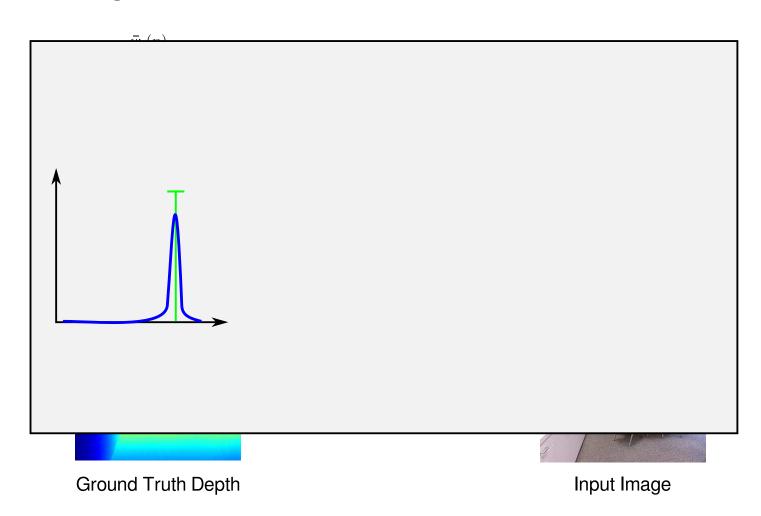
Training

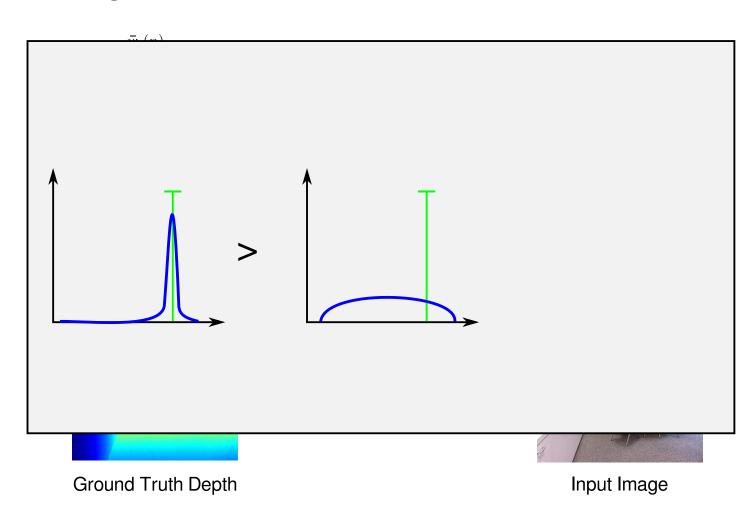


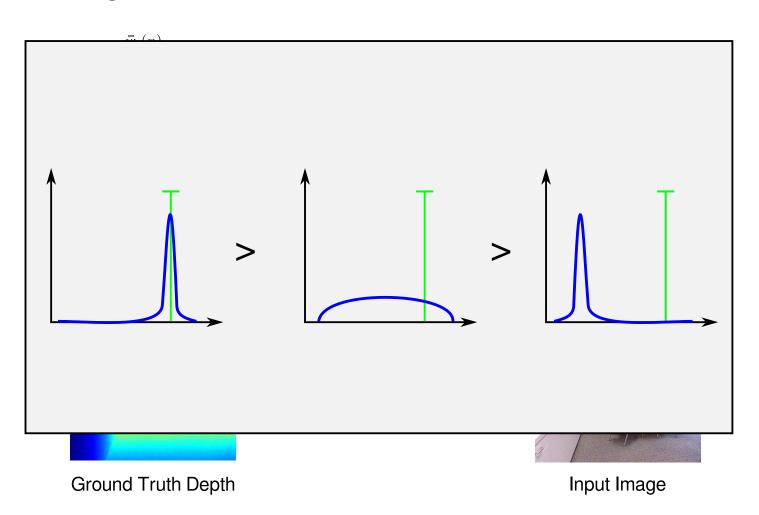
Ground Truth Depth

Input Image

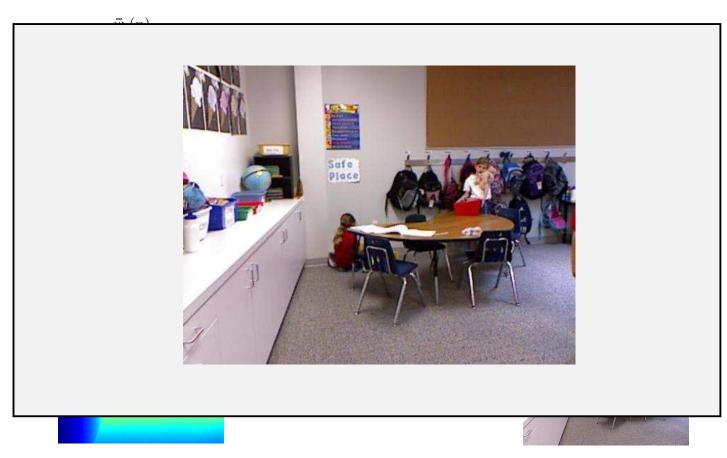






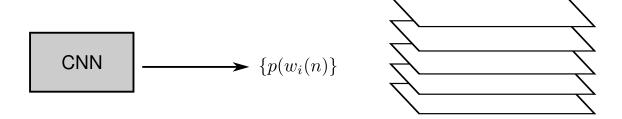


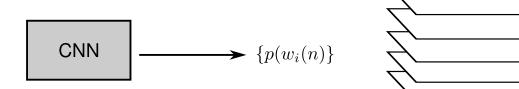
Training

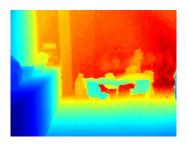


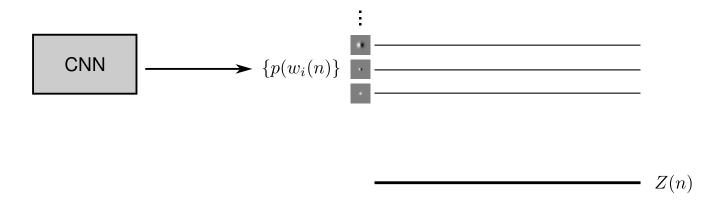
Ground Truth Depth

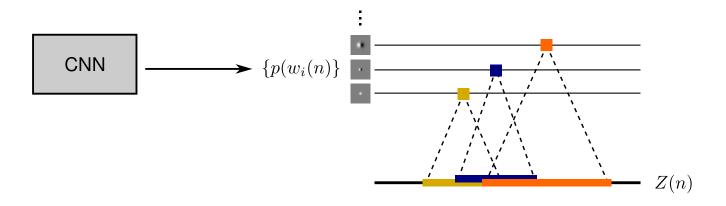
Input Image

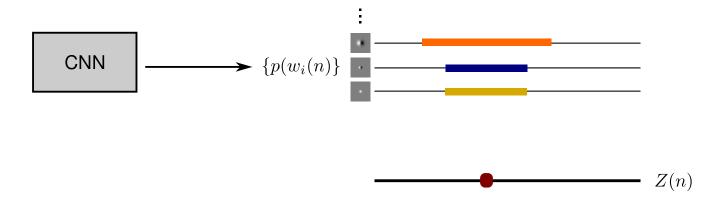


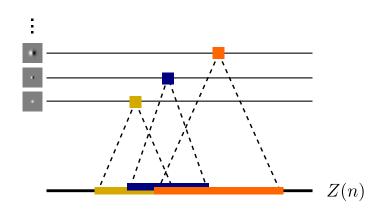






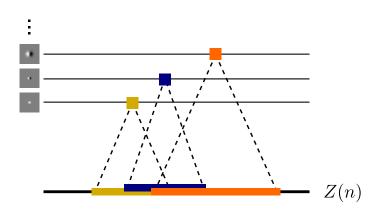






$$Z = \arg\max_{Z} \sum_{i,n} \ \log \underline{p_{i,n}} \left((Z * k_i)(n) \right)$$
 From CNN

Globalization

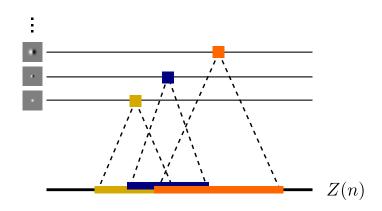


$$Z = \arg \max_{Z} \sum_{i,n} \log p_{i,n} \left((Z * k_i)(n) \right)$$

$$Z = rg \min_{Z} \min_{\left\{ \underline{w_i(n)}
ight\}} - \left[\sum_{i,n} \quad \log p_{i,n} \left(\underline{w_i(n)}
ight)
ight]$$

Auxiliary Vars for Derivatives

Globalization

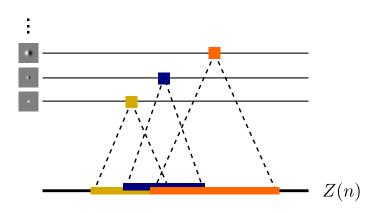


$$Z = \arg \max_{Z} \sum_{i,n} \log p_{i,n} \left((Z * k_i)(n) \right)$$

$$Z = \arg\min_{Z} \min_{\{\underline{w_i(n)}\}} - \left[\sum_{i,n} \log p_{i,n} \left(w_i(n)\right)\right] + \frac{\beta}{2} \left[\sum_{i,n} \left|w_i(n) - (Z * k_i)(n)\right|^2\right]$$

Auxiliary Vars for Derivatives

Globalization

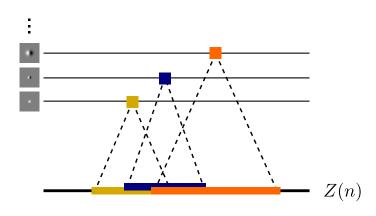


$$Z = \arg\max_{Z} \sum_{i,n} \log p_{i,n} \left((Z * k_i)(n) \right)$$

$$Z = \arg\min_{Z} \min_{\{\underline{w_i(n)}\}} - \left[\sum_{i,n} \log p_{i,n} (w_i(n)) \right] + \frac{\beta}{2} \left[\sum_{i,n} |w_i(n) - (Z * k_i)(n)|^2 \right]$$

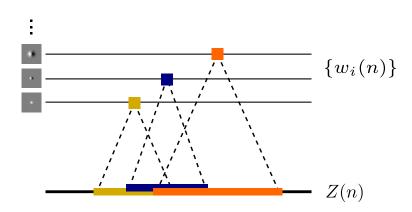
Auxiliary Vars for Derivatives

Equivalent as $\beta \to \infty$



$$Z = \arg\min_{Z} \min_{\{w_i(n)\}} - \left[\sum_{i,n} \log p_{i,n} (w_i(n)) \right] + \frac{\beta}{2} \left[\sum_{i,n} |w_i(n) - (Z * k_i)(n)|^2 \right]$$

Globalization



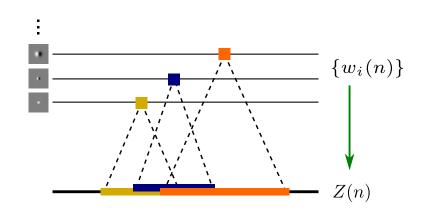
$$Z = \arg\min_{Z} \min_{\{w_i(n)\}} - \left[\sum_{i,n} \log p_{i,n} (w_i(n)) \right] + \frac{\beta}{2} \left[\sum_{i,n} |w_i(n) - (Z * k_i)(n)|^2 \right]$$

Alternatingly minimize Z and $\{w_i(n)\}$, keeping the other constant.

Globalization

Fix w, minimize wrt Z

Efficient least-squares in the Fourier-domain.

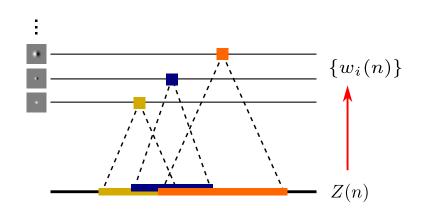


$$Z = \arg\min_{Z} \min_{\{w_i(n)\}} - \left[\sum_{i,n} \log p_{i,n} (w_i(n)) \right] + \frac{\beta}{2} \left[\sum_{i,n} |w_i(n) - (Z * k_i)(n)|^2 \right]$$

Globalization

Fix Z, minimize wrt w

Independent for each w_i(n)



$$Z = \arg\min_{Z} \min_{\{w_{i}(n)\}} - \left[\sum_{i,n} \left| \log p_{i,n} (w_{i}(n)) \right| + \frac{\beta}{2} \left[\sum_{i,n} \left| w_{i}(n) - (Z * k_{i})(n) \right|^{2} \right] \right]$$

Experimental Results

NYUv2 Depth Benchmark

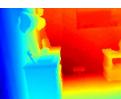
- Ground truth data from Kinect.
- 56,000 training pairs, 100 validation.
- 654 Test scenes.







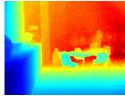


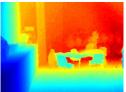


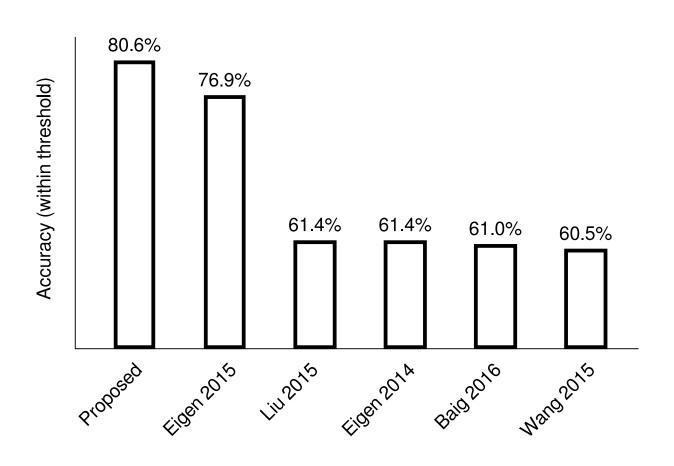


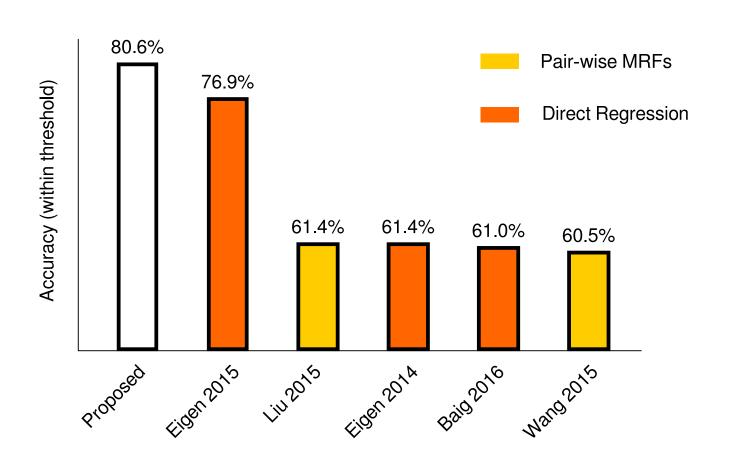


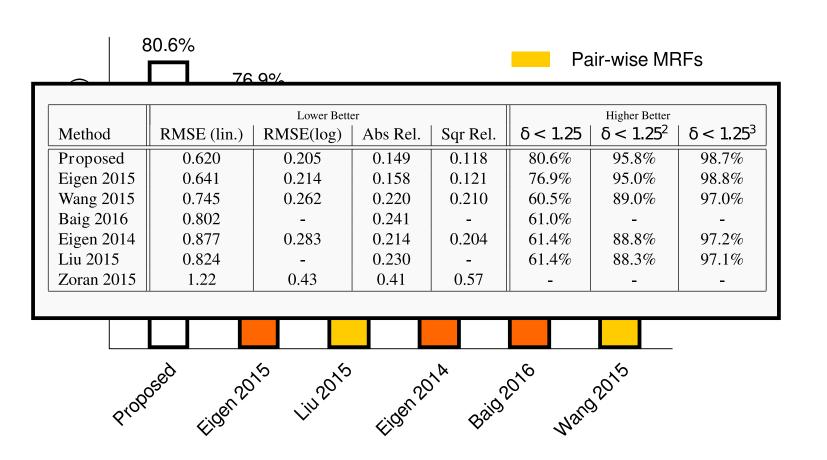






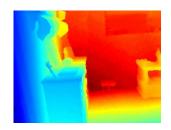








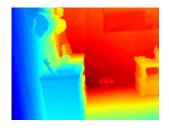
Input Image



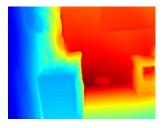
Ground Truth
Depth



Input Image



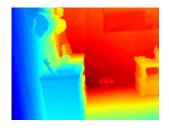
Ground Truth Depth



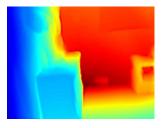
Proposed Method



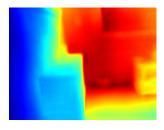
Input Image



Ground Truth
Depth



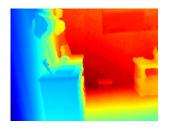
Proposed Method



Eigen 2015 (VGG)

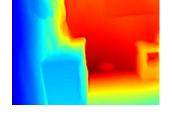


Input Image



Ground Truth Depth

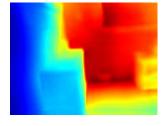




Proposed Method







Eigen 2015 (VGG)



Input Image



Ground Truth Depth

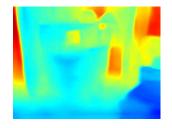




Proposed Method



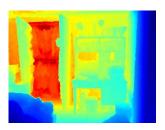




Eigen 2015 (VGG)

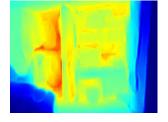


Input Image



Ground Truth Depth





Proposed Method







Eigen 2015 (VGG)

Beyond the Benchmark

$$Z = \arg \max_{Z} \sum_{i,n} \log p_{i,n} \left((Z * k_i)(n) \right)$$

Beyond the Benchmark

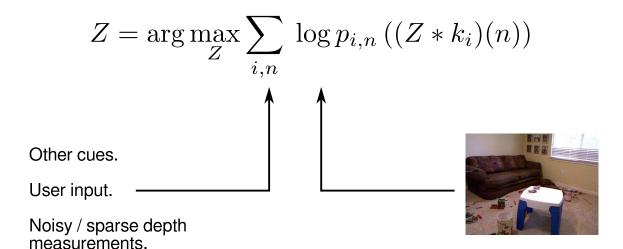
$$Z = \arg \max_{Z} \sum_{i,n} |\log p_{i,n} ((Z * k_i)(n))|$$

Rich, distributional, interpretable

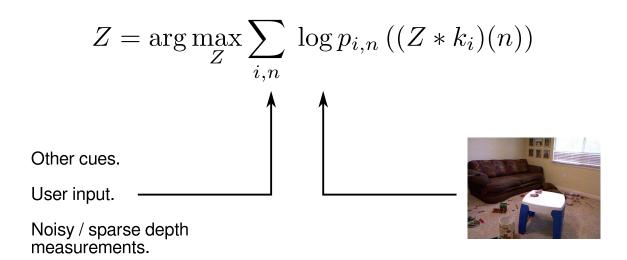
Beyond the Benchmark

$$Z = \arg\max_{Z} \sum_{i,n} \log p_{i,n} \left((Z * k_i)(n) \right)$$

Beyond the Benchmark



Beyond the Benchmark



Common substrate for local estimates from different cues.

Beyond the Benchmark

$$Z = \arg \max_{Z} \sum_{i,n} \log p_{i,n} \left((Z * k_i)(n) \right)$$

Beyond the Benchmark

$$Z = \arg \max_{Z} \sum_{i,n} \log p_{i,n} \left((Z * k_i)(n) \right)$$

$$P(Z(n) < \delta)$$

DISCUSSION

- Flavor of what a research project looks like.
- Look at group website for papers describing some of our other recent work.

Questions?